Terminator effects and dayside ionospheric current system on Pi2 pulsations

Imajo Shun¹⁾, Akimasa Yoshikawa²⁾¹⁾, Teiji Uozumi²⁾, Shinichi Ohtani³⁾, Aoi Nakamizo⁴⁾, Peter J. Chi⁵⁾

1) Department of Earth and Planetary Science, Kyushu University

- 2) International Center for Space Weather Science and Education, Kyushu University
- 3) Johns Hopkins University Applied Physics Laboratory
- 4) Corporate Planning Department, National Institute of Information and Communications Technology
- 5) Earth Planetary and Space Sciences, University of California

磁気圏起源擾乱の赤道強化現象

Pi2地磁気脈動を含む多くの磁気圏起源の擾乱 (DP2, Pc5, SCなど)は昼間側磁気赤道で振幅が増す



夜側起源のPi2がどのようにして伝わり昼間側赤道に

大きな振幅をもたらすか?

昼間側Pi2と電離圏電流

過去の研究では、昼間側地上低緯度のH成分磁場は極域電場の侵入に よる東西の電離層電流振動が原因であると解釈される。 [Kikuchi and Araki,1979, JASTP; Shinohara et al.,1998, JGR]



日射領域のPi2が電離圏電流が関与していれば<mark>朝夕昼夜境界</mark> 付近でなんらかの変化が見られるはず

Sunrise effect on Pi2 and Pc4 range wave



複数観測点での同時観測はでは未だ検証されていない。

Imajo, S., A. Yoshikawa, T. Uozumi, S. Ohtani, A. Nakamizo, R. Marshall, B. M. Shevtsov, V. A. Akulichev, U. Sukhbaatar, A. Liedloff, and K. Yumoto (2015), Pi2 pulsations observed around the dawn terminator. *J. Geophys. Res.* Space Physics, 120, 2088–2098. doi: 10.1002/2013JA019691.

朝側昼夜境界でのPi2 (21:50 UT on 16 September 2011)



- ・H成分はほぼ同位相。
- ・昼夜境界付近を境にD成分の位相が反転する。
- ・日陰、日射領域共に、南北半球でD成分の位相が反転する。

朝側昼夜境界でのPi2 (その他の例)



統計解析 (位相反転のタイミング)

朝側ASBのD成分とそのとき夜側にあるTAMのH成分の位相差を計算した

- ・D成分の位相反転のタイミングはは地方時(b)よりむしろ 日の出からの時刻(b)に従う。
- ・D成分の位相反転は100km高度の昼夜境界の約0.5時間昼間側で起こる。



夕方側昼夜境界付近のPi2

- ・日の入境界から約2-3時間昼間側でD成分の位相の 反転が見られる
- ・ 位相の反転する経度から昼間側のD成分の振幅は非
 常に小さい





昼夜境界付近のPi2の解釈



Pi2の等価電流の定義

水平成分に関して主軸に沿った変動のみに注目する

水平成分の磁場変動($\sqrt{\Delta H^2 + \Delta D^2}$)の極大時に関して等価電流の方向をマッピングする



AL index, 昼間側Pi2の波形 (23:10 UT on 2 February 2012)



Equivalent currents at the first local maximum



- The meridional component of equivalent current vectors are directed equatorward in the prenoon sector and poleward in the postnoon sector.
- The meridional component in the postnoon sector is smaller than in the prenoon sector.

Time variation of equivalent current



- ・午前午後で赤道方向に出入りするような逆成分の南北成分電流が見られる
- •午前側でより南北電流の卓越した非対称性が見られる
- ・電流系形状を維持したままPi2周期で振動している

Event 2 (2012-03-22/21:35:00)



Event 3 (2012-03-02/23:35:00)



current system which closes in the 調整電電流系の数値計算 Solid and dashed curves in 電離電源系の数値計算 demarcation lines for the reversal of the signs of the

Yoshikawa et al. [AGU fall meeting, 2012]

昼間側Pi2の等価電流はdawnとduskにFACを置いた 時の昼間側電離圏電流系とゆがみ方がよく似ている

夜側に局在したFACでも同様の昼間側電離圏電流系が出来るのか?

Ionospheric current derived by global potential solver

We solve ionospheric potential Φ produced by fixed FACs in a similar frame work of Nakamizo et al. [2012]

We derive ionospheric current $\,J\,$ from $\, E=-\nabla\cdot\Phi\,$ and $\, J=\Sigma\cdot E\,$

- Time-variable effects are negligible.
- Relative value of outputs does not depend on the magnitude of FACs.
- Conductivity tensor are calculated by the same method as Nakamizo et al. with input parameters at the case study event. This method includes the modifications by auroral precipitation [Hardy et al., 1987] and equatorial Cowling conductivity [Tsunomura, 1999].

Calculated ionospheric current distribution

- The zonal current reaches a peak at the equator because of higher conductivities modified by the Cowling effect.
- The simulated dayside current system shows the prenoon-postnoon asymmetry similar to equivalent current distributions of dayside Pi2.
- The order of the current density in the dayside low-latitude region is approximately 1/100 of the current density in the nightside auroral region.

・昼間側で閉じる電流系の非対称性に対するFACの非対称性の依存は小さい

・イベントごとのFACの構造は昼間側電流系にあまり影響しないと考えられる

Pi2のグローバル電流系による解釈

END

Thank you for your attention

Acknowledgement:

This work was supported in part by JSPS Core-to-Core Program, B. Asia-Africa Science Platforms and by Grant-in-Aid for JSPS Fellows (15J02300). MAGDAS/CPMN magnetic data were provided by the principle investigator of MAGDAS/CPMN project (http://magdas.serc.kyushu-u.ac.jp/). We thank the national institutes that support INTERMAGNET for promoting high standards of magnetic observatory practice (www.intermagnet.org). Magnetic data from KAK, CBI and, KNY are provided by JMA. We acknowledge NASA contract NAS5-02099 and V. Angelopoulos for use of data from the THEMIS Mission. U.S. Geological Survey magnetometers: Original data provided by the USGS Geomagnetism Program (http://geomag.usgs.gov). McMAC magnetometers: Peter Chi for use of the McMAC data and NSF for support through grant ATM-0245139. We acknowledge the Inter-University Upper Atmosphere Global Observation Network (IUGONET) project funded by the Ministry of Education, Culture, Sports, Science and Technology (MEXT), Japan.